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ON THE MOTION OF SEVERAL RIGID BODIES IN A VISCOUS MULTIPOLAR FLUID

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On the motion of several rigid bodies in a viscous multipolar fluid

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1 Introduction

The mathematical theory of viscous multipolar fluids, based on the general ideas of Green and Rivlin [8], was proposed by Nečas and Šilhavý [17] (see also Nečas et al. [15], [16] for relevant existence theory) in order to develop a general framework for studying viscous fluids and to present a suitable alternative to the boundary layer theory (see Bellout et al. [1]). The theory is compatible with the basic principles of thermodynamics as well as with the principle of material frame indifference. The present paper is concerned with the mathematical description of the motion of one or several rigid bodies immersed in a viscous multipolar fluid. The principal and very natural idea behind the analysis presented below is the fact that the dissipation of mechanical energy, being much stronger than for classical newtonian fluids, yields better estimates on the gradient of the velocity field, in particular, the streamlines are well-defined, which seems crucial for this class of problems partially formulated in terms of the Lagrangean coordinate system.

1.1 Bodies and motions

From the mathematical viewpoint, a *rigid body* can be identified with a connected compact subset \overline{S} of the Euclidean space R^3 , the motion of which is represented as a mapping $\eta: (0,T) \times R^3 \to R^3$, where

$$\eta(t, \cdot) : R^3 \to R^3$$
 is an isometry (1.1)

for any time $t \in (0, T)$. Throughout the whole text, we adopt the Eulerian (spatial) description of motion, where the coordinate system is attached to a fixed region of the physical space currently occupied by the fluid. The place **x** and the time $t \in (0, T)$ play the role of independent variables.

As the mappings $\eta(t, \cdot)$ are isometries, we can write

$$\eta(t, \mathbf{x}) = \mathbf{X}_g(t) + \mathbb{O}(t)(\mathbf{x} - \mathbf{X}_g(0)),$$

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where \mathbf{X}_{g} stands for the position of the center of mass at a time t, and $\mathbb{O}(t)$ is a matrix satisfying $\mathbb{O}^T \mathbb{O} = \mathbb{I}$. Assuming the motion to be absolutely continuous with respect to time, we introduce

$$\frac{\mathrm{d}}{\mathrm{d}t}\mathbf{X}_g = \mathbf{U}_g \text{ - the translation velocity}, \tag{1.2}$$

and

$$\frac{\mathrm{d}}{\mathrm{d}t}\mathbb{O}(t)\mathbb{O}^{T}(t) = \mathbb{Q}(t) \text{ - the angular velocity.}$$
(1.3)

Accordingly, the solid velocity in the Eulerian coordinate system can be written in the form

$$\mathbf{u}^{S}(t,\mathbf{x}) = \frac{\partial \eta}{\partial t}(t,\eta^{-1}(t,\mathbf{x})) = \mathbf{U}_{g}(t) + \mathbb{Q}(t)(\mathbf{x} - \mathbf{X}_{g}(t)),$$

where \mathbf{X}_g is determined through (1.2). The total force \mathbf{F}^S acting on the body \overline{S} can be written as a sum of the *body force* and the *contact force*, more specifically,

$$\mathbf{F}^{S}(t) = \int_{\overline{S}(t)} \varrho^{S} \mathbf{g}^{S} \, \mathrm{d}\mathbf{x} + \int_{\partial \overline{S}(t)} \mathbb{T}\mathbf{n} \, \mathrm{d}\sigma,$$

where \mathbb{T} denotes the Cauchy stress, \mathbf{g}^{S} is the specific body force, and

$$\overline{S}(t) = \eta(t, \overline{S}).$$

Thus Newton's second law gives rise to

$$m\frac{\mathrm{d}}{\mathrm{d}t}\mathbf{U}_g(t) = \frac{\mathrm{d}}{\mathrm{d}t}\int_{\overline{S}(t)} \varrho^S \mathbf{u}^S \,\mathrm{d}\mathbf{x} = \int_{\overline{S}(t)} \varrho^S \mathbf{g}^S \,\mathrm{d}\mathbf{x} + \int_{\partial\overline{S}(t)} \mathbb{T}\mathbf{n} \,\mathrm{d}\sigma, \qquad (1.4)$$

where m denotes the total mass of the body.

On the other hand, as the angular velocity \mathbb{Q} is skew-symmetric, there exists a vector ω such that

$$\mathbb{Q}(t)(\mathbf{x} - \mathbf{X}_g) = \omega(t) \times (\mathbf{x} - \mathbf{X}_g).$$

The balance of moment of momentum reads

$$\frac{\mathrm{d}}{\mathrm{d}t}(\mathbb{J}\omega) = \frac{\mathrm{d}}{\mathrm{d}t} \int_{\overline{S}(t)} \varrho^{S}(\mathbf{x} - \mathbf{X}_{g}) \times \mathbf{u}^{S} \,\mathrm{d}\mathbf{x} =$$

$$\int_{\partial \overline{S}(t)} (\mathbf{x} - \mathbf{X}_{g}) \times \mathbb{T}\mathbf{n} \,\mathrm{d}\sigma + \int_{\overline{S}(t)} \varrho^{S}(\mathbf{x} - \mathbf{X}_{g}) \times \mathbf{g}^{S} \,\mathrm{d}\mathbf{x},$$
(1.5)

where \mathbb{J} is the *inertial tensor* that can be identified through formula

$$\mathbb{J}\mathbf{a} \cdot \mathbf{b} = \int_{\overline{S}(t)} \varrho^{S}(\mathbf{a} \times (\mathbf{x} - \mathbf{X}_{g})) \cdot (\mathbf{b} \times (\mathbf{x} - \mathbf{X}_{g})) \, \mathrm{d}\mathbf{x}.$$

Equations (1.4), (1.5) determine completely the motion of the rigid body initially occupying the spatial domain \overline{S} .

1.2 The fluid motion

In what follows, we shall assume that the *state* of the fluid is completely determined by its density ρ^f and the velocity \mathbf{u}^f satisfying the standard mass and momentum balance equations:

$$\partial \varrho^f + \operatorname{div}_x(\varrho^f \mathbf{u}^f) = 0, \tag{1.6}$$

$$\partial_t(\varrho^f \mathbf{u}^f) + \operatorname{div}_x(\varrho^f \mathbf{u}^f \otimes \mathbf{u}^f) + \nabla_x p = \operatorname{div}_x \mathbb{S} + \varrho^f \mathbf{g}^f, \tag{1.7}$$

where the symbol p denotes the pressure, \mathbf{g}^{f} is the specific body force, and \mathbb{S} denotes the viscous stress tensor related to the total stress through Stokes' law:

$$\mathbb{T} = \mathbb{S} - p\mathbb{I}.\tag{1.8}$$

In the present paper, the effect of the temperature on the motion will be ignored. On the other hand, we consider a general *compressible* fluid so that the state equation relates the pressure to the fluid density through an empirical formula

$$p = p(\varrho^f). \tag{1.9}$$

1.3 Viscosity

The heart of the theory of multipolar fluids lies in a particular choice of constitutive equations relating the fluid stress expressed through S to the symmetric component of the velocity gradient. Very roughly indeed, one can say that, in contrast to the classical theory of newtonian fluids, the stress tensor depends on higher order gradients of the velocity field. This piece of information is sufficient in order to obtain a priori estimates yielding, in particular, strong compactness of the density ρ .

Following the seminal paper by Nečas and Šilhavý [17] we assume the viscous stress tensor S to be given as

$$\mathbb{S}[\mathbf{u}] = \sum_{n=0}^{k-1} (-1)^n \Delta^n \Big[\mu_n \Big(\nabla_x \mathbf{u} + \nabla_x \mathbf{u}^t - \frac{2}{3} \mathrm{div}_x \mathbf{u} \, \mathbb{I} \Big) + \zeta_n \mathrm{div}_x \mathbf{u} \mathbb{I} \Big], \ \mathbf{u} = \mathbf{u}^f, \quad (1.10)$$

where μ_n , ζ_n are (constant) viscosity coefficients, and the symbol Δ stands for the standard Laplace operator. Accordingly, one can speak about a k-polar fluid, the classical newtonian fluids being identified as monopolar with k = 1.

As expected, the presence of higher-order viscosities provides very strong a priori estimates on the velocity gradient, in particular, the streamlines are well defined allowing for the Lagrangean description of motion. Note that the theory of multipolar fluids requires additional "higher order" stresses to be introduced in the energy equation in order to comply with the second law of thermodynamics (see Nečas and Šilhavý [17]).

1.4 Boundary conditions

A proper choice of the boundary conditions represents one of the most delicate issues of the present theory. We adopt the hypothesis of complete adherence of the fluid to the boundaries of rigid objects yielding the *full-stick* boundary conditions

$$D_x^j \mathbf{u}^f = D_x^j \mathbf{u}^S, \ j = 0, ..., k - 1, \text{ on } \partial \overline{S}(t),$$
(1.11)

$$D_x^j \mathbf{u}^f = 0, \ j = 0, ..., k - 1, \ \text{on } \partial\Omega$$
 (1.12)

provided the flow is confined to a fixed spatial domain $\Omega \subset \mathbb{R}^3$.

Here, the symbol D_x^j denotes the vector of all spatial derivatives of order j, and conditions (1.11), (1.12) are *different* from those considered in [17], the latter being of "Neumann-type". Clearly, the boundary conditions depend on the physical properties of a given fluid and as such must be determined by experiments. In the present setting, the flowing rigid objects are supposed to be "sticky", in particular, they may be thought of as integral parts of the surrounding fluid of extremely high viscosity.

1.5 Global-in-time solutions and collisions of rigid objects

The motion of one or several rigid bodies in a viscous fluid has been a topic of numerous theoretical studies. Desjardins and Esteban [3] establish the existence of local-in-time solutions for incompressible newtonian fluids, where "local" is to be understood "up to the first collision of two rigid objects" if $\Omega \subset \mathbb{R}^2$, or "up to the blowup of the velocity gradient in a certain Sobolev norm" in the case $\Omega \subset \mathbb{R}^3$. Similar methods are proposed in [4] in order to study both compressible and incompressible case. Similar existence results "up to the first collision" were obtained by Conca et al. [2], Gunzburger et al. [9], Hoffmann and Starovoitov [12], among others.

As we have seen, the problem of existence or rather non-existence of collisions is important not only because of its practical implications, but also from the purely theoretical point of view. To the best of our knowledge, this issue remains largely open even for a two-dimensional physical domain Ω . In the 2-D case, however, there is a remarkable result by San Martin et al. [18] stating, in particular, that possible collisions, if any, must be "smooth", that means, with zero relative velocities. Another strong evidence of absence of collisions in the 2-D geometry is provided independently by Hesla [10] and Hillairet [11]. These authors show, roughly speaking, that newtonian viscosity is strong enough to prevent collisions provided the rigid objects are discs. As already pointed out, the question is completely open for a linearly viscous fluid in the realistic situation $\Omega \subset \mathbb{R}^3$ (for partial results see Starovoitov [20]).

The main objective of the present paper is to estalish the existence of globalin-time solutions for problem (1.1 - 1.12) provided $k \ge 3$. In particular, we show that collisions cannot occur in a finite time unless they were already included in the initial data. The paper is organized as follows. In Section 2, we introduce a variational (weak) formulation of the problem in the spirit of Galdi [7], Hoffmann and Starovoitov [12], Serre [19]. The main results concerning global-in-time solutions are stated in Section 3. Suitable approximate solutions are constructed in Section 4 by means of a scheme similar to that used in [6]. In particular, the method of construction is based on the idea of San Martin et al. [18], where the rigid objects are approximated by a fluid of large viscosity. Such an approach is of course intimately related to our choice of the boundary conditions specified through (1.11), (1.12). In Section 5, we derive suitable uniform estimates on the sequence of approximate solutions based on strong dissipation of the kinetic energy for multipolar fluids. In particular, the velocity field is uniformly Lipschitz continuous which facilitates the subsequent analysis considerably. Using the uniform energy estimates, we pass to the limit in the sequence of approximate solutions in order to obtain a suitable variational solution of the original problem (see Section 6).

and

2 Variational formulation

Similarly to a major part of the reference material mentioned above, our approach is based on a suitable variational formulation of the problem. In what follows, we shall assume that $\Omega \subset \mathbb{R}^3$ is a bounded domain with smooth (\mathbb{C}^∞) boundary.

2.1 Kinematics of the rigid bodies

The reference position S^i , i = 1, ..., m, of the i-th rigid body is a bounded domain in R^3 with smooth boundary. The motions are described through affine isometries $\eta^i(t, \cdot) : R^3 \to R^3, i = 1, ..., m$, that are absolutely continuous as functions of the time $t \in [0, T]$. Furthermore, we set

$$S^{i}(t) = \eta^{i}(t, S^{i}), \ i = 1, ..., m,$$

and

$$Q^{S} = \Big\{ (t, \mathbf{x}) \ \Big| \ t \in (0, T), \ \mathbf{x} \in \bigcup_{i=1}^{m} S^{i}(t) \Big\}.$$

2.2 Conservation of mass

The solid densities ρ^{S^i} as well as the fluid density ρ^f can be extended to be zero outside $S^i(t)$, and the fluid region $Q^f = ((0,T) \times \Omega) \setminus \overline{Q}^S$, respectively. In a similar way, we introduce a "global" velocity field **u**,

$$\mathbf{u}(t,\mathbf{x}) = \begin{cases} \mathbf{u}^{S^{i}} \text{ for } t \in (0,T), \ \mathbf{x} \in \overline{S}^{i}(t), \\ \mathbf{u}^{f} \text{ for } t \in (0,T), \ \mathbf{x} \in \Omega \setminus \bigcup_{i=1}^{m} \overline{S}^{i}(t), \\ 0 \text{ for } t \in (0,T), \ \mathbf{x} \in R^{3} \setminus \Omega. \end{cases}$$

The physical principle of mass conservation can be expressed through continuity equation

$$\partial_t \rho + \operatorname{div}_x(\rho \mathbf{u}) = 0 \text{ in } \mathcal{D}'((0,T) \times R^3).$$
 (2.1)

2.3 Momentum equation

For a given family of motions η^i , i = 1, ..., m, we introduce the set of admissible velocity fields

$$V_{\text{adm}}(t) = \{ \mathbf{w} \in C^k(R^3; R^3) \mid \mathbf{w} \equiv 0 \text{ on } R^3 \setminus \Omega, \ \nabla_x \mathbf{w} + \nabla_x \mathbf{w}^t \equiv 0 \text{ on } \cup_{i=1}^m \eta^i(t, S^i) \}.$$
(2.2)

Following Nečas [14], we introduce a bilinear form $((\cdot, \cdot))$ associated to the stress tensor S given by (1.10), specifically,

$$((\mathbf{v}, \mathbf{w})) = \int_{\Omega} \mathbb{S}[\mathbf{v}] : \nabla_x \mathbf{w} \, \mathrm{d}x \text{ for any } \mathbf{v}, \ \mathbf{w} \in \mathcal{D}(\Omega; R^3).$$
(2.3)

Under the natural hypothesis

$$\mu_n, \ \eta_n \ge 0, \ \mu_{k-1} > 0, \ \eta_{k-1} > 0,$$
(2.4)

it is straightforward to see that $((\cdot, \cdot))$ can be extended to a scalar product on the Sobolev space $W_0^{k,2}(\Omega; \mathbb{R}^3)$ defined as a completion of the set of compactly supported

smooth functions with respect to the norm

$$\|\mathbf{v}\|_{W_0^{k,2}(\Omega;R^3)}^2 = \sum_{n=0}^k \int_{\Omega} |D_x^n \mathbf{v}|^2 \, \mathrm{d}x$$

(see Nečas [14]).

Assuming continuity of the stresses on the boundaries of rigid objets, we can reformulate equations (1.4), (1.5), (1.7) in terms of ρ and **u** as integral identity

$$\int_{0}^{T} \int_{R^{3}} \left(\rho \mathbf{u} \cdot \partial_{t} \mathbf{w} + \rho \mathbf{u} \otimes \mathbf{u} : \nabla_{x} \mathbf{w} + p \operatorname{div}_{x} \mathbf{w} \right) \mathrm{d}\mathbf{x} \, \mathrm{d}t =$$

$$\int_{0}^{T} ((\mathbf{u}, \mathbf{w})) \, \mathrm{d}t - \int_{0}^{T} \int_{R^{3}} \rho \mathbf{g} \cdot \mathbf{w} \, \mathrm{d}\mathbf{x} \, \mathrm{d}t - \int_{R^{3}} \rho_{0} \mathbf{u}_{0} \cdot \mathbf{w}(0) \, \mathrm{d}\mathbf{x}$$

$$(2.5)$$

to be satisfied for any test function

$$\mathbf{w} \in C^1([0,T]; C^k(R^3)), \ \mathbf{w}(T) = 0, \ \mathbf{w}(t) \in V_{\text{adm}}(t) \text{ for any } t \in [0,T].$$
 (2.6)

Note that (2.5) includes the initial conditions

$$(\boldsymbol{\varrho}\mathbf{u})(0,\cdot) = \boldsymbol{\varrho}_0\mathbf{u}_0. \tag{2.7}$$

2.4 Compatibility of the "global" velocity with rigid motions

We shall say that a velocity field **u** is compatible with the family of rigid motions η^i , i = 1, ..., m provided

$$\mathbf{u}(t,\mathbf{x}) = \mathbf{u}^{S^{i}}(t,\mathbf{x}) = \frac{\partial \eta^{i}}{\partial t}(t,(\eta^{i})^{-1}(t,\mathbf{x})) \text{ for all } \mathbf{x} \in \eta^{i}(t,S^{i}), \ i = 1,...,m.$$
(2.8)

Relation (2.8) is to be satisfied for any $t \in [0, T]$.

2.5 Energy inequality

The velocity field associated to a multipolar fluid is expected regular because of the strong kinetic energy dissipation in the high velocity gradient regime. Indeed taking (formally) a test function $\mathbf{w} = -\psi(t)\mathbf{u}$, in (2.5) we obtain energy inequality

$$E(t_2) + \int_{t_1}^{t_2} ((\mathbf{u}, \mathbf{u})) \, \mathrm{d}t \le E(t_1) + \int_{t_1}^{t_2} \int_{\Omega} \rho \mathbf{g} \cdot \mathbf{u} \, \mathrm{d}x$$
(2.9)

for any $0 \le t_1 \le t_2 \le T$, with

$$E(t) = \int_{\Omega} \left(\frac{1}{2} \rho |\mathbf{u}|^2 + \rho P(\rho) \right) \, \mathrm{d}x, \qquad (2.10)$$

where P is related to the pressure p through

$$P(\varrho) = \varrho \int_{1}^{\varrho} \frac{p(z)}{z^2} \, \mathrm{d}z.$$
(2.11)

2.6 Weak (variational) solutions

Having collected all the preliminary material, we are in a position to introduce the concept of weak solution to our problem referred to hereafter as **problem** (\mathbf{P}).

Definition 2.1

Let the initial distribution of the density and the velocity field be determined through given functions ϱ_0 and \mathbf{u}_0 , respectively; the initial position of the rigid bodies being $S^i \subset \Omega$, i = 1, ..., m. We say that a family ϱ , \mathbf{u} , η^i , i = 1, ..., m, represent a variational solution of **problem (P)** on a time interval (0, T) provided the following conditions are satisfied:

• The density ρ is a non-negative bounded function, the velocity field **u** belongs to the space $L^{\infty}(0,T; L^2(\Omega; \mathbb{R}^3)) \cap L^2(0,T; W_0^{k,2}(\Omega; \mathbb{R}^3))$, and they satisfy energy inequality (2.9) for $t_1 = 0$ and a.a. $t_2 \in (0,T)$, with

$$E(0) = E_0 = \int_{\Omega} \left(\frac{1}{2} \rho_0 |\mathbf{u}_0|^2 + \rho_0 P(\rho_0) \right) \, \mathrm{d}x.$$

- We have $\rho \in C([0,T]; L^1(\Omega))$, $\rho(0) = \rho_0$, and continuity equation (2.1) holds on $(0,T) \times R^3$ provided ρ and **u** were extended to be zero outside Ω .
- Momentum equation (the integral identity) (2.5) holds for any admissible test function **w** satisfying (2.6).
- The mappings ηⁱ, i = 1,...,m are affine isometries of R³ compatible with the velocity field **u** in the sense of (2.8).

3 Global existence - main results

Our main goal is to prove the following existence result.

Theorem 3.1 Let $\Omega \subset \mathbb{R}^3$, $S^i \subset \mathbb{R}^3$, i = 1, ..., m be a family of bounded domains with boundaries of class C^{∞} such that

$$\overline{S}^{i} \cap \overline{S}^{j} = \emptyset \text{ for } i \neq j, \ \overline{S}^{i} \subset \Omega \text{ for } i = 1, ..., m.$$

$$(3.1)$$

Furthermore, assume that ρ_0 is a measurable function such that

$$0 < \varrho \le \varrho_0(x) \le \overline{\varrho} \text{ for a.a. } x \in \Omega, \tag{3.2}$$

and that

$$\mathbf{u}_0 \in W_0^{k,2}(\Omega; \mathbb{R}^3) \cap V_{\mathrm{adm}}(0) \text{ for a certain } k \ge 3.$$

$$(3.3)$$

Finally, let $\mathbf{g} \in L^{\infty}(\Omega; \mathbb{R}^3)$, and $p \in C[0, \infty)$ - a non-decreasing function be given.

Then **problem** (P) admits a variational solution ρ , \mathbf{u} , η^i , i = 1, ..., m, in the sense of Definition 2.1 on an arbitrary time interval (0,T). Moreover, we have

$$\overline{S}^{i}(t) \cap \overline{S}^{j}(t) = \emptyset \text{ for } i \neq j, \ \overline{S}^{i}(t) \subset \Omega \text{ for } i = 1, ..., m$$
(3.4)

for any $t \in [0,T]$, that means, the motion is smooth without any collision of two or several rigid objects in a finite time.

As already pointed out several times, the main ingredient of the proof are strong dissipation estimates resulting from energy inequality (2.9). Since $k \ge 3$, the velocity field **u** is a priori bounded in the space $L^2(0,T; C_0^1(\overline{\Omega}))$, in particular, the streamlines (characteristics) can be identified with the unique solution of the system

$$\frac{\mathrm{d}}{\mathrm{d}t}\mathbf{X}(t,\mathbf{x}) = \mathbf{u}(t,\mathbf{X}(t,\mathbf{x})), \ t > 0, \ \mathbf{X}(0,\mathbf{x}) = \mathbf{x}.$$
(3.5)

Accordingly, the (unique) weak solution ρ of (2.1) satisfying $\rho(0) = \rho_0$ is given by formula

$$\varrho(t, \mathbf{X}(t, \mathbf{x})) = \varrho_0(x) \exp\left(-\int_0^t \operatorname{div}_x \mathbf{u}(s, \mathbf{X}(s, \mathbf{x})) \, \mathrm{d}s\right), \ \mathbf{x} \in \mathbb{R}^3.$$
(3.6)

Note that, in the absence of collisions, one can deduce from (2.5) that

$$\partial_t(\varrho \mathbf{u}) \in L^2(0, T; W^{-k,2}(\Omega; \mathbb{R}^3)), \tag{3.7}$$

which, combined with (2.9), gives rise to

$$\mathbf{u} \in C([0,T]; L^2(\Omega; \mathbb{R}^3)).$$
 (3.8)

In particular, formula (3.6) makes sense for any weak solution in the sense of Definition 2.1.

The rest of the paper is devoted to the proof of Theorem 3.1.

4 Approximate problems

Let $\{\mathbf{v}_n\}_{n=1}^{\infty} \subset \mathcal{D}(\Omega; \mathbb{R}^3)$ be a basis of the Hilbert space $W_0^{k,2}(\Omega; \mathbb{R}^3)$. Following San Martin et al. [18] we introduce the approximate **problem** (**P**)_{*n*, ε}:

$$\int_{0}^{T} \int_{R^{3}} \left(\varrho_{n,\varepsilon} \mathbf{u}_{n,\varepsilon} \cdot \partial_{t} \mathbf{w} + \varrho_{n,\varepsilon} \mathbf{u}_{n,\varepsilon} \otimes \mathbf{u}_{n,\varepsilon} : \nabla_{x} \mathbf{w} + p(\varrho_{n,\varepsilon}) \operatorname{div}_{x} \mathbf{w} \right) \mathrm{d}\mathbf{x} \, \mathrm{d}t = \quad (4.1)$$
$$\int_{0}^{T} \int_{\Omega} M_{\varepsilon}(\chi_{n,\varepsilon}) [\nabla_{x} \mathbf{u}_{n,\varepsilon} + \nabla_{x}^{t} \mathbf{u}_{n,\varepsilon}] : [\nabla_{x} \mathbf{w} + \nabla_{x}^{t} \mathbf{w}] \, \mathrm{d}x \, \mathrm{d}t + \\\int_{0}^{T} ((\mathbf{u}_{n,\varepsilon}, \mathbf{w})) \, \mathrm{d}t - \int_{0}^{T} \int_{R^{3}} \varrho_{n,\varepsilon} \mathbf{g} \cdot \mathbf{w} \, \mathrm{d}\mathbf{x} \, \mathrm{d}t - \int_{R^{3}} \varrho_{0} \mathbf{u}_{0} \cdot \mathbf{w}(0) \, \mathrm{d}\mathbf{x}$$
he satisfied for any test function

to be satisfied for any test function

$$\mathbf{w} \in C^1([0,T];X_n), \ \mathbf{w}(T) = 0, X_n = \operatorname{span}\{\mathbf{v}_1, ..., \mathbf{v}_n\}.$$
 (4.2)

Here, $\rho_{n,\varepsilon}$ is determined through formula

$$\varrho_{n,\varepsilon}(t, \mathbf{X}_{n,\varepsilon}(t, \mathbf{x})) = \varrho_0(x) \exp\left(-\int_0^t \operatorname{div}_x \mathbf{u}_{n,\varepsilon}(s, \mathbf{X}_{n,\varepsilon}(s, \mathbf{x})) \, \mathrm{d}s\right), \ \mathbf{x} \in \mathbb{R}^3$$
(4.3)

while

$$\chi_{n,\varepsilon}(t, \mathbf{X}_{n,\varepsilon}(t, \mathbf{x})) = \chi_{0,\varepsilon}(\mathbf{x}) \ge 0, \qquad (4.4)$$

where

$$\frac{\mathrm{d}}{\mathrm{d}t}\mathbf{X}_{n,\varepsilon}(t,\mathbf{x}) = \mathbf{u}_{n,\varepsilon}(t,\mathbf{X}_{n,\varepsilon}(t,\mathbf{x})), \ t > 0, \ \mathbf{X}_{n,\varepsilon}(0,\mathbf{x}) = \mathbf{x}.$$
(4.5)

Note that $\chi_{n,\varepsilon}$ is the unique distributional solution of the equation

$$\partial_t \chi_{n,\varepsilon} + \mathbf{u}_{n,\varepsilon} \cdot \nabla_x \chi_{n,\varepsilon} = 0, \ \chi_{n,\varepsilon}(0) = \chi_{0,\varepsilon}$$

(cf. DiPerna and Lions [5]).

The functions M_{ε} belong to the class $C^1[0,\infty) \cap BC[0,\infty)$ for any fixed $\varepsilon > 0$.

Problem $(\mathbf{P})_{n,\varepsilon}$ can be solved by means of the standard fixed-point argument used in [13, Section 2], the presence of the additional viscosity coefficient M_{ε} requiring only minor modifications.

5 Uniform estimates

Similarly to Section 2.5, one can show that the approximate solutions $\rho_{n,\varepsilon}$, $\mathbf{u}_{n,\varepsilon}$ satisfy energy equality

$$E_{n,\varepsilon}(t_2) + \int_{t_1}^{t_2} \int_{\Omega} M_{\varepsilon}(\chi_{n,\varepsilon}) |\nabla_x \mathbf{u}_{n,\varepsilon} + \nabla_x^t \mathbf{u}_{n,\varepsilon}|^2 \, \mathrm{d}x \, \mathrm{d}t + \int_{t_1}^{t_2} \left((\mathbf{u}_{n,\varepsilon}, \mathbf{u}_{n,\varepsilon}) \right) \, \mathrm{d}t = (5.1)$$
$$E_{n,\varepsilon}(t_1) + \int_{t_1}^{t_2} \int_{\Omega} \varrho_{n,\varepsilon} \mathbf{g} \cdot \mathbf{u}_{n,\varepsilon} \, \mathrm{d}x$$

for any $0 \le t_1 \le t_2 \le T$, with

$$E_{n,\varepsilon}(t) = \int_{\Omega} \left(\frac{1}{2} \varrho_{n,\varepsilon} |\mathbf{u}_{n,\varepsilon}|^2 + \varrho_{n,\varepsilon} P(\varrho_{n,\varepsilon}) \right) \,\mathrm{d}x,\tag{5.2}$$

$$E_{n,\varepsilon}(0) = E_0 = \int_{\Omega} \left(\frac{1}{2} \varrho_0 |\mathbf{u}_0|^2 + \varrho_0 P(\varrho_0) \right) \,\mathrm{d}x.$$
(5.3)

Under the hypotheses of Theorem 3.1, it is a routine matter to check that (5.1 - 5.3) give rise to uniform estimates:

$$\{\mathbf{u}_{n,\varepsilon}\}_{n,\varepsilon} \text{ bounded in } L^2(0,T;W_0^{k,2}(\Omega;R^3)),$$
(5.4)

$$\underline{\varrho} \exp\left(-T \|\operatorname{div}_{x} \mathbf{u}_{n,\varepsilon}\|_{L^{1}(0,T;L^{\infty}(\Omega))}\right) \leq \varrho_{n,\varepsilon}(t,\mathbf{x}) \leq \overline{\varrho} \exp\left(T \|\operatorname{div}_{x} \mathbf{u}_{n,\varepsilon}\|_{L^{1}(0,T;L^{\infty}(\Omega))}\right)$$
(5.5)

for a.a. $t \in (0, T)$, $\mathbf{x} \in \Omega$, and

$$\{\mathbf{u}_{n,\varepsilon}\}_{n,\varepsilon}$$
 bounded in $L^{\infty}(0,T;L^2(\Omega;R^3)).$ (5.6)

In addition,

$$\int_{0}^{T} \int_{\Omega} M_{\varepsilon}(\chi_{n,\varepsilon}) |\nabla_{x} \mathbf{u}_{n,\varepsilon} + \nabla_{x}^{t} \mathbf{u}_{n,\varepsilon}|^{2} \, \mathrm{d}x \, \mathrm{d}t \le \mathrm{const},$$
(5.7)

where the bound is uniform with respect to n, ε .

6 Convergence

6.1 The limit $n \to \infty$

With $\varepsilon > 0$ fixed, our aim is to let $n \to \infty$ in the family of approximate solutions $\rho_{n,\varepsilon}$, $\mathbf{u}_{n,\varepsilon}$ constructed in Section 4. To begin with, estimates (5.4), (5.6) yield

$$\mathbf{u}_{n,\varepsilon} \to \mathbf{u}_{\varepsilon} \text{ weakly in } L^2(0,T; W_0^{k,2}(\Omega; R^3)) \text{ and weakly-}(*) \text{ in } L^\infty(0,T; L^2(\Omega; R^3))$$
(6.1)

for $n \to \infty$, at least for a suitable subsequence.

Similarly, by virtue of (4.3), (5.5), we can assume that

$$\varrho_{n,\varepsilon} \to \varrho_{\varepsilon} \text{ in } C_{\text{weak}}([0,T]; L^1(\Omega)) \text{ and weakly-}(*) \text{ in } L^\infty(0,T; L^\infty(\Omega)).$$
(6.2)

Consequently, combining (6.1), (6.2), we conclude that

$$\varrho_{n,\varepsilon} \mathbf{u}_{n,\varepsilon} \to \varrho_{\varepsilon} \mathbf{u}_{\varepsilon} \text{ weakly-}(^*) \text{ in } L^{\infty}(0,T;L^2(\Omega;R^3)).$$
(6.3)

Moreover, it follows from (4.1) that

$$\{t \mapsto \int_{\Omega} (\varrho_{n,\varepsilon} \mathbf{u}_{n,\varepsilon})(t) \cdot \mathbf{v}_n \, \mathrm{d}x\} \to \{t \mapsto \int_{\Omega} (\varrho_{\varepsilon} \mathbf{u}_{\varepsilon})(t) \cdot \mathbf{v}_n \, \mathrm{d}x\} \text{ in } C[0,T] \text{ for } n = 1, 2, ...;$$

whence

$$\varrho_{n,\varepsilon} \mathbf{u}_{n,\varepsilon} \to \varrho_{\varepsilon} \mathbf{u}_{\varepsilon} \text{ in } C_{\text{weak}}([0,T]; L^2(\Omega; R^3)).$$
(6.4)

Relations (6.1), (6.4) imply

$$\varrho_{n,\varepsilon} \mathbf{u}_{n,\varepsilon} \cdot \mathbf{u}_{n,\varepsilon} \to \varrho_{\varepsilon} \mathbf{u}_{\varepsilon} \cdot \mathbf{u}_{\varepsilon} \text{ weakly in } L^2((0,T) \times \Omega); \tag{6.5}$$

therefore, in view of (4.3),

$$\mathbf{u}_{n,\varepsilon} \to \mathbf{u}_{\varepsilon} \text{ in } L^2(0,T;L^2(\Omega;R^3)).$$
 (6.6)

Consequently, (6.1), (6.6) and a simple interpolation argument yield

$$\mathbf{u}_{n,\varepsilon} \to \mathbf{u}_{\varepsilon} \text{ in } L^2(0,T; W_0^{k-1,2}(\Omega; R^3)),$$
(6.7)

in particular,

$$\operatorname{div}_{x}\mathbf{u}_{n,\varepsilon} \to \operatorname{div}_{x}\mathbf{u}_{\varepsilon} \text{ in } L^{2}(0,T;W_{0}^{k-2,2}(\Omega;R^{3})).$$
(6.8)

Since $\rho_{n,\varepsilon}$ satisfy continuity equation (2.1), one can deduce from (6.2), (6.8) that

$$\varrho_{n,\varepsilon} \to \varrho_{\varepsilon} \text{ in } L^1((0,T) \times \Omega).$$
(6.9)

In a similar way, one can show

$$\chi_{n,\varepsilon} \to \chi_{\varepsilon} \text{ in } L^1((0,T) \times \Omega).$$
 (6.10)

Thus we have shown there are functions ρ_{ε} , \mathbf{u}_{ε} such that

$$\int_{0}^{T} \int_{R^{3}} \left(\varrho_{\varepsilon} \mathbf{u}_{\varepsilon} \cdot \partial_{t} \mathbf{w} + \varrho_{\varepsilon} \mathbf{u}_{\varepsilon} \otimes \mathbf{u}_{\varepsilon} : \nabla_{x} \mathbf{w} + p(\varrho_{\varepsilon}) \operatorname{div}_{x} \mathbf{w} \right) \mathrm{d}\mathbf{x} \, \mathrm{d}t = \qquad (6.11)$$
$$\int_{0}^{T} \int_{\Omega} M_{\varepsilon}(\chi_{\varepsilon}) [\nabla_{x} \mathbf{u}_{\varepsilon} + \nabla_{x}^{t} \mathbf{u}_{\varepsilon}] : [\nabla_{x} \mathbf{w} + \nabla_{x}^{t} \mathbf{w}] \, \mathrm{d}x \, \mathrm{d}t + \qquad (6.11)$$

$$\int_0^T ((\mathbf{u}_{\varepsilon}, \mathbf{w})) \, \mathrm{d}t - \int_0^T \int_{R^3} \varrho_{\varepsilon} \mathbf{g} \cdot \mathbf{w} \, \mathrm{d}\mathbf{x} \, \mathrm{d}t - \int_{R^3} \varrho_0 \mathbf{u}_0 \cdot \mathbf{w}(0) \, \mathrm{d}\mathbf{x}$$

to be satisfied for any test function

$$\mathbf{w} \in C^1([0,T]; W_0^{k,2}(\Omega, R^3)), \ \mathbf{w}(T) = 0.$$
 (6.12)

Furthermore, we have

$$\varrho_{\varepsilon}(t, \mathbf{X}_{\varepsilon}(t, \mathbf{x})) = \varrho_{0}(x) \exp\left(-\int_{0}^{t} \operatorname{div}_{x} \mathbf{u}_{\varepsilon}(s, \mathbf{X}_{\varepsilon}(s, \mathbf{x})) \, \mathrm{d}s\right), \ \mathbf{x} \in \mathbb{R}^{3}$$
(6.13)

and

$$\chi_{\varepsilon}(t, \mathbf{X}_{\varepsilon}(t, \mathbf{x})) = \chi_{0, \varepsilon}(\mathbf{x}) \ge 0, \qquad (6.14)$$

where

$$\frac{\mathrm{d}}{\mathrm{d}t}\mathbf{X}_{\varepsilon}(t,\mathbf{x}) = \mathbf{u}_{\varepsilon}(t,\mathbf{X}_{\varepsilon}(t,\mathbf{x})), \ t > 0, \ \mathbf{X}_{\varepsilon}(0) = \mathbf{x}.$$
(6.15)

In addition, the energy inequality

$$E_{\varepsilon}(t_{2}) + \int_{t_{1}}^{t_{2}} \int_{\Omega} M_{\varepsilon}(\chi_{\varepsilon}) |\nabla_{x} \mathbf{u}_{\varepsilon} + \nabla_{x}^{t} \mathbf{u}_{\varepsilon}|^{2} \, \mathrm{d}x \, \mathrm{d}t + \int_{t_{1}}^{t_{2}} ((\mathbf{u}_{\varepsilon}, \mathbf{u}_{\varepsilon})) \, \mathrm{d}t \leq \qquad (6.16)$$

$$E_{\varepsilon}(t_{1}) + \int_{t_{1}}^{t_{2}} \int_{\Omega} \varrho_{\varepsilon} \mathbf{g} \cdot \mathbf{u}_{\varepsilon} \, \mathrm{d}x$$

$$f_{\varepsilon}(t_{1}) = \int_{t_{1}}^{t_{2}} \int_{\Omega} \varrho_{\varepsilon} \mathbf{g} \cdot \mathbf{u}_{\varepsilon} \, \mathrm{d}x$$

holds for any $0 \le t_1 \le T$ and a.a. $t_2 \in (t_1, T)$, where

$$E_{\varepsilon}(t) = \int_{\Omega} \left(\frac{1}{2} \varrho_{\varepsilon} |\mathbf{u}_{\varepsilon}|^2 + \varrho_{\varepsilon} P(\varrho_{\varepsilon}) \right) \, \mathrm{d}x, \ E_{\varepsilon}(0) = E_0 = \int_{\Omega} \left(\frac{1}{2} \varrho_0 |\mathbf{u}_0|^2 + \varrho_0 P(\varrho_0) \right) \, \mathrm{d}x.$$

6.2 The limit for $\varepsilon \to 0$

Adopting the idea of San Martin et al. [18] we take

$$M_{\varepsilon}(z) = \frac{1}{\varepsilon} \max\{z, 0\}, \qquad (6.17)$$

$$\chi_{\varepsilon,0} = \chi_0 \in C^1(\overline{\Omega}), \ \chi_0(\mathbf{x}) = \begin{cases} 0 \text{ for } \mathbf{x} \in \overline{\Omega} \setminus \bigcup_{i=1}^m \overline{S}^i, \\ > 0 \text{ for } \mathbf{x} \in \bigcup_{i=1}^m S^i. \end{cases}$$
(6.18)

Our ultimate goal is to let $\varepsilon \to 0$ in (6.11 - 6.16) in order to recover the global-in-time solution of **problem (P)**, the existence of which is claimed in Theorem 3.1.

To begin with, we can assume, by virtue of (6.16), that

$$\mathbf{u}_{\varepsilon} \to \mathbf{u}$$
 weakly in $L^2(0, T; W_0^{k,2}(\Omega; \mathbb{R}^3))$ (6.19)

passing to a suitable subsequence as the case may be. Moreover, as $\{\varrho_{\varepsilon}\}_{\varepsilon>0}$ is bounded below away from zero in view of (6.13), we have

$$\mathbf{u}_{\varepsilon} \to \mathbf{u} \text{ weakly-}(^*) \text{ in } L^{\infty}(0,T;L^2(\Omega;R^3)).$$
 (6.20)

As stated in (6.13), $\{\varrho_{\varepsilon}\}_{\varepsilon>0}$ solve (in the sense of distributions) continuity equation (2.1) supplemented with the initial datum ϱ_0 ; whence we have

$$\varrho_{\varepsilon} \to \rho \text{ in, say, } C_{\text{weak}}([0,T]; L^2(\Omega)),$$
(6.21)

$$\varrho_{\varepsilon} \mathbf{u}_{\varepsilon} \to \varrho \mathbf{u} \text{ weakly-}(^*) \text{ in } L^{\infty}(0,T;L^2(\Omega;R^3)).$$
(6.22)

Thus we are allowed to conclude that ρ , \mathbf{u} , extended to be zero outside Ω , solve equation (2.1). In addition, ρ is uniquely determined by ρ_0 and the velocity field \mathbf{u} , and $\rho \in C([0,T]; L^1(\Omega))$ (cf. DiPerna and Lions [5]).

In order to identify the family of isometries η^i , i = 1, ..., m, we need the following auxiliary result proved in [6, Proposition 5.1].

Lemma 6.1 Let $\mathbf{u}_{\varepsilon} = \mathbf{u}_{\varepsilon}(t, \mathbf{x})$ be a family of Caratheodory functions such that

$$\int_0^T \|\mathbf{u}_{\varepsilon}(t)\|_{W^{1,\infty}(R^3;R^3)}^2 \, \mathrm{d}t \le \text{const}$$

uniformly with respect to $\varepsilon \to 0$. Let $S_{\varepsilon} \subset \mathbb{R}^3$ be a family of open sets such that

$$\mathbf{db}[S_{\varepsilon}] \to \mathbf{db}[S] \text{ in } C_{\mathrm{loc}}(R^3),$$

where $S \subset \mathbb{R}^3$ is an open set and the symbol **db** denotes the signed distance from the boundary:

$$\mathbf{db}[S](\mathbf{x}) = \operatorname{dist}[\mathbf{x}, \overline{R^3 \setminus S}] - \operatorname{dist}[\mathbf{x}, \overline{S}], \ \operatorname{dist}[\mathbf{x}, K] \equiv \min_{\mathbf{y} \in K} |\mathbf{x} - \mathbf{y}|.$$

Denote by \mathbf{X}_{ε} the unique solution of the problem

$$\frac{\mathrm{d}}{\mathrm{d}t} \mathbf{X}_{\varepsilon}(t, \mathbf{x}) = \mathbf{u}_{\varepsilon}(t, \mathbf{X}_{\varepsilon}(t, \mathbf{x})), \ 0 < t < T, \ \mathbf{X}_{\varepsilon}(0, \mathbf{x}) = \mathbf{x}.$$
(6.23)

Then, extracting a suitable subsequence if necessary, we have

$$\mathbf{u}_{\varepsilon} \to \mathbf{u} \ weakly (*) \ in \ L^2(0,T; W^{1,\infty}(\mathbb{R}^3;\mathbb{R}^3)),$$

and

$$\mathbf{X}_{\varepsilon}(t,\cdot) \to \mathbf{X}(t,\cdot)$$
 in $C_{\mathrm{loc}}(R^3)$ uniformly in $t \in [0,T]$,

where \mathbf{X} solves

$$\frac{\mathrm{d}}{\mathrm{d}t} \mathbf{X}(t, \mathbf{x}) = \mathbf{u}(t, \mathbf{X}(t, \mathbf{x})), \ 0 < t < T, \ \mathbf{X}(0, \mathbf{x}) = \mathbf{x}.$$
(6.24)

Moreover,

$$\mathbf{db}[S_{\varepsilon}(t)] \to \mathbf{db}[S(t)] \text{ in } C_{\mathrm{loc}}(R^3) \text{ uniformly in } t \in [0,T],$$

where we have ser

$$S_{\varepsilon}(t) = \mathbf{X}_{\varepsilon}(t, S_{\varepsilon}), \ S(t) = \mathbf{X}(t, S).$$

Consider the domains $S^{i}(t)$, i = 1, ..., m occupied by the images of the rigid bodies S^{i} , i = 1, ..., m under the flow induced by (6.24):

$$S^{i}(t) = \mathbf{X}(t, S^{i}), \ i = 1, ..., m.$$

Since the velocity field **u** belongs to the class $L^2(0,T; C^1(\mathbb{R}^3; \mathbb{R}^3))$, we have, in accordance with hypothesis (3.1),

$$\overline{S}^{i}(t) \cap \overline{S}^{j}(t) = \emptyset \text{ for } i \neq j, \ \overline{S}^{i} \subset \Omega \text{ for } i = 1, ..., m$$
(6.25)

for all $t \in [0, T]$, that is to say, there is no collision of two or more "rigid" objects. Let

$$(t,\mathbf{x}) \in \bigcup_{t \in (0,T)} S^i(t)$$

In accordance with Lemma 6.1, there is a small open neighbourhood V of (t, \mathbf{x}) such that

$$V \subset \overline{V} \subset \bigcup_{t \in (0,T)} \mathbf{X}_{\varepsilon}(t, S^i)$$
 for all $\varepsilon > 0$ small enough,

where \mathbf{X}_{ε} are determined through (6.15).

Consequently, combining energy inequality (6.16) with (6.14), (6.17), (6.18) we conclude that

$$\nabla_x \mathbf{u} + \nabla_x^t \mathbf{u} = 0$$
 a.a. on V.

As the point (t, \mathbf{x}) was arbitrary, we have

$$\nabla_x \mathbf{u} + \nabla_x^t \mathbf{u} = 0 \text{ a.a. on } \cup_{i=1}^m \cup_{t \in (0,T)} S^i(t).$$
(6.26)

It is a routine matter to deduce from (6.24), (6.26) that

$$\mathbf{X}(t, \cdot) : S^i \to R^3$$
 is an isometry for any $i = 1, ..., m$ and any fixed $t \in [0, T]$;

whence we can set

$$\eta^{i}(t, \cdot) : R^{3} \to R^{3}, \ \eta^{i}(t, \mathbf{x}) = \mathbf{X}(t, \mathbf{x}) \text{ for all } \mathbf{x} \in S^{i}, \ t \in [0, T], \ i = 1, ..., m$$
 (6.27)

Clearly, the family η^i , i = 1, ..., m is compatible with the vector field **u** in the sense of Definition 2.1.

In order to complete the proof of Theorem 3.1, we have to pass to the limit for $\varepsilon \to 0$ in (6.11) to recover momentum equation (2.5). To this end, first observe that

 $\{\varrho_{\varepsilon}\mathbf{u}_{\varepsilon}\}_{\varepsilon>0}$ is precompact in $C_{\text{weak}}([t_1, t_2], L^2(B; \mathbb{R}^3))$ (6.28)

whenever $([t_1, t_2] \times \overline{B}) \cap \bigcup_{i=1}^m \bigcup_{t \in [0,T]} \overline{S}^i(t) = \emptyset.$

In particular, in accordance with (6.20),

$$\varrho_{\varepsilon} \mathbf{u}_{\varepsilon} \otimes \mathbf{u}_{\varepsilon} \to \mathbb{Q} \text{ weakly in } L^2(0,T;L^2(\Omega;R^{3\times 3})),$$
(6.29)

where

$$\mathbb{Q} = \rho \mathbf{u} \otimes \mathbf{u} \text{ on } \Omega \setminus \bigcup_{i=1}^{m} \bigcup_{t \in (0,T)} S^{i}(t).$$
(6.30)

As a byproduct of (6.30) we get

$$\mathbf{u}_{\varepsilon} \to \mathbf{u} \text{ in } L^2 \Big(\Omega \setminus \bigcup_{i=1}^m \bigcup_{t \in (0,T)} S^i(t); R^3 \Big);$$
(6.31)

whence, having used the estimates on $\operatorname{div}_x \mathbf{u}_{\varepsilon}$ resulting from (6.16), we conclude that

$$\operatorname{div}_{x} \mathbf{u}_{\varepsilon} \to \operatorname{div}_{x} \mathbf{u} \text{ in } L^{1}((0,T) \times \Omega).$$
(6.32)

Since ρ_{ε} solve the continuity equation, relation (6.32) implies

$$\varrho_{\varepsilon} \to \varrho \text{ in } L^1((0,T) \times \Omega).$$
(6.33)

Using (6.29), (6.33) one can let $\varepsilon \to 0$ in (6.11) in order to recover (2.5) at least for any test function **w** such that

$$\mathbf{w} \in C^1([0,T]; C^k(R^3)), \ \mathbf{w}(T) = 0, \ \mathbf{w}(t) \in V^{\delta}_{\mathrm{adm}}(t) \text{ for any } t \in [0,T],$$

where

$$V_{\text{adm}}^{\delta}(t) = \{ \mathbf{w} \in C^k(\mathbb{R}^3; \mathbb{R}^3) \mid \mathbf{w} \equiv 0 \text{ on a } \delta \text{-neighbourhood of } \overline{\mathbb{R}^3 \setminus \Omega_2} \}$$

$$\nabla_x \mathbf{w} + \nabla_x \mathbf{w}^t \equiv 0$$
 on a δ -neighbourhood of $\bigcup_{i=1}^m \overline{\eta^i(t, S^i)}$, $\delta > 0$.

Note that

$$\int_0^T \int_\Omega \varrho_\varepsilon \mathbf{u}_\varepsilon \otimes \mathbf{u}_\varepsilon : \nabla_x \mathbf{w} \, \mathrm{d}x \, \mathrm{d}t = \frac{1}{2} \int_0^T \int_\Omega \varrho_\varepsilon \mathbf{u}_\varepsilon \otimes \mathbf{u}_\varepsilon : (\nabla_x \mathbf{w} + \nabla_x^t \mathbf{w}) \, \mathrm{d}x \, \mathrm{d}t.$$

Since there is no contact of rigid objets (see (6.25)), it is easy to extend validity of (2.5) to all test functions in the class (2.6) via density argument.

Finally, by virtue of (6.19), (6.33), energy inequality (2.9) holds for $t_1 = 0$ and a.a. $t_2 \in (0, T)$ as required in Definition 2.1. Theorem 3.1 has been proved.

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